

# Quasiclassical Solutions of the Schrödinger Equations as a Consequence of the Nonlinear Problem

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## Abstract

Quasiclassic method of solving of the Schrödinger equation with quadratic Hamiltonian is used to derive solutions of Klein-Fock equation for the particle in the constant magnetic field and the jumping magnetic field.

Let us consider the relativistic wave equation

$$\begin{aligned} (D_\lambda^2 - k_0^2) \psi &= 0, \\ k_0 &= \frac{mc}{\hbar}, \\ D_\mu &= \frac{\partial}{\partial x_\mu} + \frac{ie}{\hbar c} A_\mu \end{aligned} \tag{1}$$

with the electromagnetic field of the form

$$\begin{aligned} A_\mu = A_\mu^T &= \frac{\mathbf{H}}{2}(-x_2, x_1, 0, 0), \\ \mathbf{H} &= \text{const.} \end{aligned} \tag{2}$$

It should be noted that the Maxwell equations let to use the magnetic field  $\mathbf{H}$  depending on the variable  $\xi = ct - z$  (or  $\eta = ct + z$ ). We provide the following investigation of Eq.(1) in the characteristics representation

$$\begin{aligned} \xi &= ct - z, \\ \eta &= ct + z, \\ \vec{x}_\perp &= (x_1, x_2, 0) \end{aligned} \tag{3}$$

which may be introduced in covariant form. We introduce the axial and transverse coordinates

$$\begin{aligned} x &= x_1 + ix_2, \\ x^+ &= x_1 - ix_2. \end{aligned} \tag{4}$$

Thus the equation (1) has the form

$$\left( D_{\perp}^2 - 4 \frac{\partial^2}{\partial \xi \partial \eta} - k_0^2 \right) \psi = 0 \quad (5)$$

where

$$D_{\perp}^2 = 4 \frac{\partial^2}{\partial x \partial x^+} + \frac{e\mathbf{H}}{2\hbar c} \left( x^+ \frac{\partial}{\partial x^+} - x \frac{\partial}{\partial x} \right) - \frac{e^2 \mathbf{H}^2}{16\hbar^2 c^2} x^+ x. \quad (6)$$

If the operator  $D_{\perp}^2$  does not depend on the variables  $\xi, \eta$  we can apply to Eq.(5) the Fourier-Bessel transformation

$$\psi(x^T, \xi, \eta) = \int J_0(\sqrt{\xi\eta}Q) \varphi(x, x^+, Q) Q dQ \quad (7)$$

and to consider the Goursat problem solutions [1]. If the operator  $D_{\perp}^2$  depends on the characteristics variable (that is  $\mathbf{H} = \mathbf{H}(\xi)$ ) then we apply to Eq.(5) the Fourier transformation in respect to the advanced relativistic time  $\eta$  [2]

$$\psi(x^T, \xi, \eta) = \frac{1}{2\pi\hbar} \int d\beta \exp\left(\frac{-i\beta}{\hbar}\eta\right) \varphi(x, x^+, \beta) \quad (8)$$

and then we obtain "nonrelativistic" equation for function  $\varphi$

$$\left( \frac{\hbar}{i} \frac{\partial}{\partial \xi} + \mathbf{H} \right) \varphi = 0,$$

$$\mathbf{H} = -\frac{\hbar^2}{4\beta} \left( D_{\perp}^2 - k_0^2 \right). \quad (9)$$

To find the quasiclassical solutions of Eq.(8) we introduce

$$\varphi(x, x^+, \beta, \xi) = \phi(\xi, \beta) \exp\left(\frac{i}{\hbar} S(x, x^+, \xi, \beta)\right) \quad (10)$$

with the complex amplitude  $\phi$  and action  $S$ . Two equations arise

$$\begin{aligned} \frac{\partial \phi}{\partial \xi} + \frac{1}{\beta} \cdot \frac{\partial^2 S}{\partial x \partial x^+} \phi &= 0, \\ \frac{\partial S}{\partial \xi} + \frac{1}{\beta} \frac{\partial S}{\partial x} \frac{\partial S}{\partial x^+} - \frac{i\Omega}{2\beta} \left( x^+ \frac{\partial S}{\partial x^+} - x \frac{\partial S}{\partial x} \right) &= 0 \end{aligned} \quad (11)$$

where  $\Omega = \frac{eH}{\hbar c}$ . The Hamiltonian (8) is quadratic and so it can be supposed that the action  $S$  has the form

$$S = A(\xi, \beta) + B(\xi, \beta)x + B^+(\xi, \beta)x^+ + C(\xi, \beta)x^+x \quad (12)$$

and then the problem reduces to the solution of the system of ordinary differential equations of first order

$$\frac{dA}{d\xi} + \frac{1}{\beta} B^+ B + \frac{m^2 c^2}{4\beta} = 0, \quad (13)$$

$$\frac{dB}{d\xi} + \frac{1}{\beta} \left( C + \frac{i\Omega}{2} \right) B = 0, \quad \frac{dB^+}{d\xi} + \frac{1}{\beta} \left( C - \frac{i\Omega}{2} B^+ \right) = 0, \quad (14)$$

$$\frac{dC}{d\xi} + \frac{1}{\beta} \left( C^2 + \frac{\Omega^2}{4} \right) = 0. \quad (15)$$

The nonlinear equation (15) is principal and we call the function  $C$  key-function. It may be shown that all other coefficients in the action  $S$  and amplitude  $\phi$  can be expressed via the key-function  $C$ .

The Eq.(15) and its solutions have the following properties ( $\Omega$  is a constant):

1. If  $C$  is a solution then  $\tilde{C} = -\frac{\Omega^2}{4} \frac{1}{C}$  is a solution also.

2. A function of  $C$  and  $C^+$  is a solution  $C(1 + \Omega) = C(-\Omega)$ .

3. There is a pure imaginary solution if  $C = C_0 = \pm \frac{i\Omega}{2}$ .

The solution of Eq.(15) can be found by linearization of Eq.(15) by means of applying the Cayley–Klein transformation of the complex function  $C$

$$F = \frac{C - i\Omega/2}{C + i\Omega/2},$$

$$C = \frac{i\Omega}{2} \cdot \frac{1 + F}{1 - F} \quad (16)$$

where  $F$  must be the solution of the linear equation

$$\frac{dF}{d\xi} + \frac{i\Omega}{\beta} F = 0 \quad (17)$$

with the periodical solutions. That gives the key-function

$$C(\xi) = \frac{i\Omega}{2} \cdot \frac{1 + b \cdot \exp\left(\frac{-i\Omega}{\beta\xi}\right)}{1 - b \cdot \exp\left(\frac{-i\Omega}{\beta\xi}\right)} \quad (18)$$

where  $b$  is a complex parameter of the constant of integration. It follows from Eq.(18) that the imaginary part of the key-function has a form

$$\text{Im } C = \frac{\Omega(1 - \rho^2)}{2 \left( 1 + \rho^2 - 2\rho \cos\left(\frac{\Omega\xi}{2\beta} - \vartheta\right) \right)} \quad (19)$$

where  $\rho = b^*b$  and  $\vartheta = \arctan\left(\frac{\text{Im } b}{\text{Re } b}\right)$ . The importance of the parameter  $b$  is seen from the investigation of the behaviour of the function  $\text{Im } C$ . If  $\rho^2 < 1$  and  $\text{Im } C > 0$  then the states are normalizable. Passing to the region where  $\rho^2 > 1$  it is seen that function  $\text{Im } C$  disappears if  $\rho^2 = 1$  and then the sign changes of the Eq.(19).

There is the periodicity of key-function  $C$  in relativistic retarded time due to its interesting property  $\left(T = \frac{2\pi\beta}{\Omega}\right)$

$$\langle C(\xi) \rangle = \frac{1}{T} \int_{\xi_0}^{\xi_0+T} d\xi' C(\xi') = \frac{i\Omega}{2}. \quad (20)$$

The value  $C_0 = \frac{i\Omega}{2}$  corresponds to the stable coherent state we obtain some singular state when  $b = 1$  and  $\text{Im } C = 0$ ,  $\text{Re } C = \frac{\Omega}{2} \cot \frac{\Omega\xi}{2\beta}$ . This solution represents the fundamental solution of the problem and it may be used as the Green function. The region  $\rho^2 < 1$  represents the pulsating states unknown in traditional theory. The problem to be developed in this approach is the behaviour of the system by the jumps of magnetic field  $H$  and the jumps of the value  $\Omega$ . The continuity of solutions ( $\varphi$  or  $\psi$ ) is guaranteed by the continuity of the key-function  $C$ , and when the frequency  $\Omega_1$  changes at  $\xi = \xi_0$  into  $\Omega_2$  the condition of continuity is

$$C(\Omega_1, \xi_0) = C(\Omega_2, \xi_0).$$

We remind that if  $\Omega = 0$  then the solution for a free particle is

$$C(\xi) = \frac{C_0}{1 + C_0\xi/\beta}. \tag{21}$$

It two particular cases can be considered:

1. the case of the plane wave if  $C_0 = 0$ .
2. the case of the free Green function of the Schrödinger equation if  $C_0 \rightarrow \infty$ , then the key-function  $C$  is  $C = \frac{\beta}{\xi}$  and

$$\varphi = \frac{\phi_0}{\xi} \exp\left(\frac{i\beta x_1^2}{\hbar\xi} - \frac{im^2 c^2 \xi}{4\hbar\beta}\right). \tag{22}$$

We emphasize that if the magnetic field  $H \rightarrow 0$  then the state with  $b = 1$  and  $C = \frac{\beta}{\xi}$  is the Green function and there is a stable coherent state if  $b = 0$ . If the parameter  $0 < b < 1$  then the stable coherent state with  $C = C_0 = \frac{i\Omega}{2}$  become pulsately.

Let before the point  $\xi_0$  the frequency was  $\Omega_0$  and after  $\xi_0$  become it  $\Omega_1$ . Then the condition of continuity gives ( $b_0 = 0$ )

$$b_1 = \frac{\Omega_0 - \Omega_1}{\Omega_0 + \Omega_1} \exp\left(\frac{i\Omega_1\xi_0}{\beta}\right), \tag{23}$$

and

$$b_1^* b_1 = \left(\frac{\Omega_0 - \Omega_1}{\Omega_0 + \Omega_1}\right)^2 < 1.$$

The inverse process of transfer from pulsating state with given  $b_1$  into stable coherent state gives

$$b_1 = \frac{\Omega_1 - \Omega_0}{\Omega_0 + \Omega_1} \exp\left(\frac{i\Omega_0\xi_0}{\beta}\right), \tag{24}$$

and  $b_1^* b_1 < 1$  once again.

Making two such the jumps ( $\xi_1 < \xi_0$ ) we obtain the condition of conservation of the stable coherent state

$$\exp\left(\frac{i\Omega_1(\xi_1 - \xi_0)}{\beta}\right) = \frac{\Omega_0 - \Omega_1}{\Omega_0 + \Omega_1} \cdot \frac{\Omega_1 + \Omega_2}{\Omega_2 - \Omega_1} \tag{25}$$

that gives the restriction on the time interval  $\Delta\xi = \xi_1 - \xi_0$ . This is possible at  $\Omega_1 = \Omega_0$  only. In this case Eq.(25) has the solution

$$\Delta\xi_n = \frac{\pi\beta}{2\Omega_1}(2n + 1). \quad (26)$$

After the rectangle impulse of amplitude of  $\Omega_1 - \Omega_0$  and the duration of  $\Delta\xi_n$  the system returns into the stable coherent state.

## References

- [1] Courant R., Partial Differential Equations, New York, London, 1962.
- [2] Borghardt A., Karpenko D., On the class of the integral transformations in relativistic quantum theory, Kiev, 1975, Preprint /ITP-75-19E, 16p.