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# CONSERVATION LAWS FOR THE SCHRÖDINGER–NEWTON EQUATIONS

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In this Letter a first-order Lagrangian for the Schrödinger–Newton equations is derived by modifying a second-order Lagrangian proposed by Christian [Exactly soluble sector of quantum gravity, *Phys. Rev. D* **56**(8) (1997) 4844–4877]. Then Noether's theorem is applied to the Lie point symmetries determined by Robertshaw and Tod [Lie point symmetries and an approximate solution for the Schrödinger–Newton equations, *Nonlinearity* **19**(7) (2006) 1507–1514] in order to find conservation laws of the Schrödinger–Newton equations.

Keywords: Schrödinger-Newton equations; calculus of variations; Noether's theorem.

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# 1. Introduction

The Schrödinger–Newton equations consist of a system of partial differential equations introduced by Penrose [9], and in dimensionless units<sup>a</sup> they are<sup>b</sup>:

$$i\psi_t = -\Delta\psi + \phi\psi, \qquad (1.2a)$$

$$\Delta \phi = |\psi|^2, \tag{1.2b}$$

where  $\Delta$  is the Laplacian in  $\mathbb{R}^3(x, y, z)$ . The Lie point symmetry algebra admitted by the Schrödinger–Newton equations was determined by Robertshaw and Tod [10]. We rewrite

$$-\mathrm{i}\psi_t^* = -\Delta\psi^* + \phi\psi^*. \tag{1.1}$$

<sup>&</sup>lt;sup>a</sup>For more details see Harrison [4].

<sup>&</sup>lt;sup>b</sup>Including the equation satisfied by the complex conjugate  $\psi^*$  of the wave function  $\psi$ , i.e.

the Schrödinger–Newton equations as follows:

$$iu_t = -\Delta u + uw, \tag{1.3a}$$

$$-iv_t = -\Delta v + vw, \tag{1.3b}$$

$$\Delta w = uv. \tag{1.3c}$$

With  $\psi = u$ ,  $\psi^* = v$  and  $\phi = w$ . The Lie point symmetries found in [10] are:

• scaling:

$$\boldsymbol{v}_1 = 2t\partial_t + x\partial_x + y\partial_y + z\partial_z - 2u\partial_u - 2v\partial_v - 2w\partial_w, \qquad (1.4)$$

• spatial rotations:

$$\boldsymbol{v}_2 = y\partial_x - x\partial_y,\tag{1.5a}$$

$$\boldsymbol{v}_3 = z\partial_x - x\partial_z,\tag{1.5b}$$

$$\boldsymbol{v}_4 = z\partial_y - y\partial_z,\tag{1.5c}$$

• translation in time and in all spatial directions:

$$\boldsymbol{v}_5 = \partial_t, \tag{1.6a}$$

$$\boldsymbol{v}_6 = \partial_x, \tag{1.6b}$$

$$\boldsymbol{v}_7 = \partial_y, \tag{1.6c}$$

$$\boldsymbol{v}_8 = \partial_z, \tag{1.6d}$$

• phase change in the wave function:

$$\boldsymbol{v}_{9}(\Omega) = \mathrm{i}\Omega(t)(u\partial_{u} - v\partial_{v}) - \Omega'(t)\partial_{w}, \qquad (1.7)$$

• generalized Galilean group:

$$\boldsymbol{v}_{10}(a_1) = a_1(t)\partial_x + \frac{i}{2}a_1'(t)(u\partial_u - v\partial_v) - \frac{1}{2}a_1''(t)\partial_w, \qquad (1.8a)$$

$$\mathbf{v}_{11}(a_2) = a_2(t)\partial_y + \frac{i}{2}a'_2(t)(u\partial_u - v\partial_v) - \frac{1}{2}a''_2(t)\partial_w,$$
 (1.8b)

$$\mathbf{v}_{12}(a_3) = a_3(t)\partial_z + \frac{i}{2}a'_3(t)(u\partial_u - v\partial_v) - \frac{1}{2}a''_3(t)\partial_w.$$
 (1.8c)

# 2. First-Order Variational Formulation

Several variational formulations were proposed for the Schrödinger–Newton equations. At first they were recovered in the frame of the field theory by Kibble and Randjbar-Daemi [5] using an Hamiltonian formalism. Other variational formulations were given by Christian [2] and Diósi [3]. In his Ph.D. thesis Harrison [4], following Tod [11], proposed to derive the Schrödinger–Newton equations (1.3) by means of the functional:

$$S_1[u,v] = \int_{t_0}^{+\infty} \iiint_{\mathbb{R}^3} \left[ (\operatorname{grad} u, \operatorname{grad} v) + \frac{1}{2}wuv + \frac{\mathrm{i}}{2}(uv_t - vu_t) \right] \mathrm{d}^3 \mathbf{x} \mathrm{d}t,$$
(2.1)

with the condition  $\Delta w = uv$  implied. Following Tod [11] he later proposed to solve the equation for w with the usual method of the Green function in  $\mathbb{R}^3$ :

$$w(\boldsymbol{x},t) = \frac{-1}{4\pi} \iiint_{\mathbb{R}^3} \frac{u(\boldsymbol{y},t)v(\boldsymbol{y},t)}{|\boldsymbol{x}-\boldsymbol{y}|} \mathrm{d}^3 \boldsymbol{y}.$$
 (2.2)

However this leads to the functional:

$$S_{2}[u,v] = \int_{t_{0}}^{+\infty} \iiint_{\mathbb{R}^{3}} \left[ (\operatorname{grad} u, \operatorname{grad} v) - uv \frac{1}{8\pi} \iiint_{\mathbb{R}^{3}} \frac{u(\boldsymbol{y}, t)v(\boldsymbol{y}, t)}{|\boldsymbol{x} - \boldsymbol{y}|} \mathrm{d}^{3}\boldsymbol{y} \right] \mathrm{d}^{3}\boldsymbol{x} \mathrm{d}t + \int_{t_{0}}^{+\infty} \iiint_{\mathbb{R}^{3}} \left[ \frac{\mathrm{i}}{2} (uv_{t} - vu_{t}) \right] \mathrm{d}^{3}\boldsymbol{x} \mathrm{d}t,$$

$$(2.3)$$

that is nonlocal, and therefore the usual rules of Calculus of Variation do not hold [1].

Instead we take into consideration a paper by Christian [2] where the Schrödinger– Newton equations were derived from the following variational principle:

$$\hat{S}[u,v,w] = \int_{t_0}^{+\infty} \iiint_{\mathbb{R}^3} \left[ \frac{1}{2} w \Delta w - (\text{grad } u, \text{grad } v) - \frac{\mathrm{i}}{2} (uv_t - u_t v) - wuv \right] \mathrm{d}^3 \boldsymbol{x} \mathrm{d}t. \quad (2.4)$$

It was obtained by matching the Newton–Cartan theory and Quantum Mechanics. The functional (2.4) is not a first-order functional, but we note that using the general formula:

$$f\Delta g = \operatorname{Div}(f \operatorname{grad} g) - (\operatorname{grad} f, \operatorname{grad} g), \quad f, g \in \mathcal{C}^2(\mathbb{R}^3, \mathbb{R})$$
(2.5)

we may write:

$$w\Delta w = \operatorname{Div}(w \operatorname{grad} w) - |\operatorname{grad} w|^2, \qquad (2.6)$$

and apply Gauss theorem to the integral of the first term in (2.6) in order to get:

$$\iiint_{\mathbb{R}^3} \operatorname{Div}(w \operatorname{grad} w) \mathrm{d}^3 \boldsymbol{x} = \lim_{R \to \infty} \iint_{S^2(R)} (w \operatorname{grad} w, \boldsymbol{n}_{S^2(R)}) \mathrm{d}^2 \boldsymbol{q}, \tag{2.7}$$

with  $S^2(R) = \{ x \in \mathbb{R}^3 \mid |x| = R \}$ . This integral may be put equal to zero by assuming that the function w behaves like 1/r in a neighborhood of infinity if the usual boundary condition for the Poisson equation holds.

Thus we obtain<sup>c</sup> the following first-order variational principle:

$$S[u, v, w] = \int_{t_0}^{+\infty} \iiint_{\mathbb{R}^3} \left[ \frac{1}{2} |\operatorname{grad} w|^2 + wuv \right] \mathrm{d}^3 \boldsymbol{x} \mathrm{d}t + \int_{t_0}^{+\infty} \iiint_{\mathbb{R}^3} \left[ (\operatorname{grad} u, \operatorname{grad} v) + \frac{\mathrm{i}}{2} (uv_t - u_t v) \right] \mathrm{d}^3 \boldsymbol{x} \mathrm{d}t.$$
(2.8)

It is easy to verify that (2.8) indeed yields the Schrödinger–Newton equations (1.3).

 $^{\rm c}{\rm Up}$  to a multiplicative minus sign that does not effect the Euler–Lagrange equations.

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#### 3. Conservation Laws

Given a first-order variational principle in p independent variables  $\boldsymbol{x} = (x^1, \ldots, x^p)$  and q dependent variables  $\boldsymbol{u} = (u^1, \ldots, u^q)$  in some functional space of functions defined over a connected set A with smooth boundary:

$$S[\boldsymbol{u}] = \int \cdots \int_{A} L(\boldsymbol{x}, \boldsymbol{u}^{(1)}) \mathrm{d}^{p} \boldsymbol{x}$$
(3.1)

a transformation group with infinitesimal generator given by a vector field  $\boldsymbol{v}$  leaves the variational principle unchanged iff we can find a *p*-tuple  $\boldsymbol{B}(\boldsymbol{x}, \boldsymbol{u}^{(1)}) = (B_1(\boldsymbol{x}, \boldsymbol{u}^{(1)}), \ldots, B_p(\boldsymbol{x}, \boldsymbol{u}^{(1)}))$  such that [8, Chap. 4]:

$$\operatorname{pr}^{(1)}\boldsymbol{v}(L) + L\operatorname{Div}\boldsymbol{\xi} = \operatorname{Div}\boldsymbol{B}.$$
(3.2)

Where  $\boldsymbol{\xi} = (\xi^1, \dots, \xi^p)$ . In that case we call  $\boldsymbol{v}$  a variational symmetry<sup>d</sup> for the variational principle (3.1).

We look for variational symmetries among the Lie point symmetries of the corresponding Euler–Lagrange equations.

A conservation law for a system of differential equations  $\Theta(x, u^{(n)}) = 0$  is a divergence expression vanishing identically along the solutions of the system:

$$\operatorname{Div} \boldsymbol{P} = 0. \tag{3.3}$$

The Schrödinger–Newton equations are a dynamical system such that their conservation laws can be put in the form [8]:

$$D_t T + \operatorname{Div} \boldsymbol{K} = 0 \tag{3.4}$$

along the solution of the system. The scalar  $T(\boldsymbol{x}, t, \boldsymbol{u}^{(1)})$  is the conserved density and the vector  $\boldsymbol{K}(\boldsymbol{x}, t, \boldsymbol{u}^{(1)})$  the associated flux.

A general theorem contained in Olver's book [8] tells us that if  $K(x, t, u(x)^{(n)}) \to 0$  for  $x \to \partial A$  then

$$\int \cdots \int_{A} T(\boldsymbol{x}, t, \boldsymbol{u}(\boldsymbol{x})^{(n)}) \mathrm{d}^{p} \boldsymbol{x} = \text{constant.}$$
(3.5)

A corollary to Noether's theorem to be found in her famous 1918-paper [6] states that given a variational symmetry v there is an explicit formula for the vector P in (3.3), i.e.:

$$P_i = \sum_{\alpha=1}^q \eta_\alpha \frac{\partial L}{\partial u_i^\alpha} + \xi^i L - B_i - \sum_{\alpha=1}^q \sum_{j=1}^p \xi^j u_j^\alpha \frac{\partial L}{\partial u_i^\alpha}.$$
(3.6)

We wrote an *ad hoc* REDUCE interactive program that is based upon that by Nucci for finding Lie symmetries [7]. It verifies the condition (3.2) and then returns the conserved density and the associated flux given by Eq. (3.6). We found out that with respect to the

<sup>&</sup>lt;sup>d</sup>In [8] Olver calls a vector field v a variational symmetry if  $B \equiv 0$  and a divergence variational symmetry if  $B \neq 0$ . Since variational symmetries are a particular class of divergence variational symmetries we prefer to call the latter just variational symmetries.

functional (2.8) all the Lie point symmetries of the Schrödinger–Newton equations, except  $v_1$  as expected, are variational symmetries.

Introducing the shorthand notation:

$$\mathcal{E} = uvw + (\text{grad } u, \text{grad } v) + \frac{1}{2}|\text{grad } w|^2, \qquad (3.7a)$$

$$\Pi^{i} = \frac{i}{2}(u_{i}v - uv_{i}), \quad i = x, y, z, t,$$
(3.7b)

$$\Phi^{ij} = u_i v_j + u_j v_i + w_i w_j, \quad i, j = x, y, z, t,$$
(3.7c)

$$\Lambda^{ij} = x^j \Pi^i - x^i \Pi^j, \quad i, j = x, y, z, \tag{3.7d}$$

$$\varepsilon^{i} = \mathcal{E} - 2u_{i}v_{i} - w_{i}^{2}, \quad i = x, y, z,$$
(3.7e)

we find that:

- $v_1$  is not a variational symmetry because the condition (3.2) is not satisfied,
- $v_2$  is a variational symmetry with  $B_2 = 0$  and

$$T_2 = \Lambda^{xy},\tag{3.8a}$$

$$\boldsymbol{K}_2 = (y(-\Pi^t + \varepsilon^x) + x\Phi^{xy}, x(\Pi^t - \varepsilon^y) - y\Phi^{xy}, x\Phi^{yz} - y\Phi^{xz}), \quad (3.8b)$$

•  $v_3$  is a variational symmetry with  $B_3 = 0$  and

$$T_3 = \Lambda^{xz},\tag{3.9a}$$

$$\boldsymbol{K}_3 = (z(\varepsilon^x - \Pi^t) + x\Phi^{xz}, x\Phi^{yz} - z\Phi^{xy}, x(\Pi^t - \varepsilon^z) - z\Phi^{xz}), \quad (3.9b)$$

•  $v_4$  is a variational symmetry with  $B_4 = 0$  and

$$T_4 = \Lambda^{zy},\tag{3.10a}$$

$$\boldsymbol{K}_4 = (y\Phi^{xz} - z\Phi^{xy}, z(\varepsilon^y - \Pi^t) + y\Phi^{xy}, y(\Pi^t - \varepsilon^z) - z\Phi^{yz}), \quad (3.10b)$$

•  $\boldsymbol{v}_5$  is a variational symmetry with  $\boldsymbol{B}_5 = \boldsymbol{0}$  and

$$T_5 = \mathcal{E},\tag{3.11a}$$

$$\mathbf{K}_5 = (-\Phi^{tx}, -\Phi^{ty}, -\Phi^{tz}),$$
 (3.11b)

•  $\boldsymbol{v}_6$  is a variational symmetry with  $\boldsymbol{B}_6 = \boldsymbol{0}$  and

$$T_6 = \Pi^x, \tag{3.12a}$$

$$\boldsymbol{K}_6 = (\varepsilon^x - \Pi^t, -\Phi^{xy}, -\Phi^{xz}), \qquad (3.12b)$$

•  $v_7$  is a variational symmetry with  $B_7 = 0$  and

$$T_7 = \Pi^y, \tag{3.13a}$$

$$\boldsymbol{K}_7 = (-\Phi^{xy}, \varepsilon^y - \Pi^t, -\Phi^{yz}), \qquad (3.13b)$$

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- $\boldsymbol{v}_8$  is a variational symmetry with  $\boldsymbol{B}_8 = \boldsymbol{0}$  and

$$T_8 = \Pi^z, \tag{3.14a}$$

$$\boldsymbol{K}_8 = (-\Phi^{xz}, -\Phi^{yz}, \varepsilon^z - \Pi^t), \qquad (3.14b)$$

•  $v_9$  is a variational symmetry with  $B_9 = 0$  and

$$T_9 = \Omega(t)uv, \tag{3.15a}$$

$$\boldsymbol{K}_{9} = \begin{pmatrix} 2\Omega(t)\Pi^{x} - \Omega'(t)w_{x} \\ 2\Omega(t)\Pi^{y} - \Omega'(t)w_{y} \\ 2\Omega(t)\Pi^{z} - \Omega'(t)w_{z} \end{pmatrix}^{t}, \qquad (3.15b)$$

•  $\boldsymbol{v}_{10}$  is a variational symmetry with  $\boldsymbol{B}_{10} = (0, -a_1''(t)w/2, 0, 0)$  and

$$T_{10} = a_1(t)\Pi^x + \frac{1}{2}a'_1(t)uvx, \qquad (3.16a)$$
$$K_{10} = \begin{pmatrix} a_1(t)(\varepsilon^x - \Pi^t) - a'_1(t)x\Pi^x + \frac{1}{2}a''_1(t)(w - w_x x) \\ -a_1(t)\Phi^{xy} - a'_1(t)x\Pi^y - \frac{1}{2}a''_1(t)w_y x \\ -a_1(t)\Phi^{xz} - a'_1(t)x\Pi^z - \frac{1}{2}a''_1(t)w_z x \end{pmatrix}^t, \qquad (3.16b)$$

•  $v_{11}$  is a variational symmetry with  $B_{11} = (0, 0, -a_2''(t)w/2, 0)$  and

$$T_{11} = a_2(t)\Pi^y + \frac{1}{2}a'_2(t)uvy, \qquad (3.17a)$$
$$K_{11} = \begin{pmatrix} -a_2(t)\Phi^{xy} - a'_2(t)y\Pi^x - \frac{1}{2}a''_2(t)w_xy \\ a_2(t)(\varepsilon^y - \Pi^t) - a'_2(t)y\Pi^y + \frac{1}{2}a''_2(t)(w - w_yy) \\ -a_2(t)\Phi^{yz} - a'_2(t)y\Pi^z - \frac{1}{2}a''_1(t)w_zy \end{pmatrix}^t, \qquad (3.17b)$$

•  $\boldsymbol{v}_{12}$  is a variational symmetry with  $\boldsymbol{B}_{12}=(0,0,0,-a_3''(t)w/2)$  and

$$T_{12} = a_3(t)\Pi^z + \frac{1}{2}a'_3(t)uvz, \qquad (3.18a)$$
$$K_{12} = \begin{pmatrix} -a_3(t)\Phi^{xz} - a'_3(t)z\Pi^x - \frac{1}{2}a''_3(t)w_xz, \\ -a_3(t)\Phi^{yz} - a'_3(t)y\Pi^y - \frac{1}{2}a''_3(t)w_yz, \\ a_3(t)(\varepsilon^z - \Pi^t) - a'_3(t)z\Pi^z + \frac{1}{2}a''_3(t)(w - w_zz) \end{pmatrix}^t. \qquad (3.18b)$$

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Assuming the boundary conditions  $u, v, w \xrightarrow{|x| \to \infty} 0$  then Eq. (3.5) yields the following conserved quantities of the Schrödinger–Newton equations:

• energy:

$$E = \iiint_{\mathbb{R}^3} \mathcal{E} \mathrm{d}^3 \boldsymbol{x}, \tag{3.19}$$

• angular momenta:

$$L^{ij} = \iiint_{\mathbb{R}^3} \Lambda^{ij} \mathrm{d}^3 \boldsymbol{x}, \quad i, j = x, y, z,$$
(3.20)

• linear momenta:

$$p^{i} = \iiint_{\mathbb{R}^{3}} \Pi^{i} \mathrm{d}^{3} \boldsymbol{x}, \quad i = x, y, z,$$
(3.21)

• generalized probability:

$$P_g(\Omega(t)) = \iiint_{\mathbb{R}^3} \Omega(t) uv d^3 \boldsymbol{x}$$
(3.22)

that becomes the usual probability if  $\Omega(t) = 1$ , i.e.

$$P = \iiint_{\mathbb{R}^3} uv \mathrm{d}^3 \boldsymbol{x},\tag{3.23}$$

• generalized linear momenta:

$$h^{i}(a_{i}(t)) = \iiint_{\mathbb{R}^{3}} \left( a_{i}(t)\Pi^{i} + \frac{1}{2}a_{i}'(t)uvx^{i} \right) \mathrm{d}^{3}\boldsymbol{x}, \quad i = x, y, z.$$
(3.24)

The conservation of energy<sup>e</sup> (3.19), the angular (3.20) and linear momenta (3.21) and the usual probability (3.23) were recovered by Harrison [4] but without any consideration of symmetries. As far as we know the generalized probability (3.22) and generalized linear momenta (3.24) are new conserved quantities of the Schrödinger–Newton equations.

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